

## VI Identical particles

In principle QM formalism easy to generalize to  $N$  particles. E.g. coordinates  $\vec{r}_i$  and spins  $(\vec{s}_i^2, s_{iz})$  form a complete commuting set of observables, and wave function would be

$$\Psi(\underbrace{\vec{r}_1, s_{1z}, s_{1z}^2; \vec{r}_2, s_{2z}, s_{2z}^2; \dots}_{\text{shorthand } s_i}; \vec{r}_N, s_N, s_N^2) = \langle \vec{r}_1, s_1, \vec{r}_2, s_2, \dots, \vec{r}_N, s_N | \Psi \rangle$$

If particles are independent, i.e. mutually noninteracting  
 $\rightarrow$  separate variables  
 $\rightarrow$  wave function product of individual wave functions

$$\Psi = \varphi_1(\vec{r}_1, s_1) \varphi_2(\vec{r}_2, s_2) \dots \varphi_N(\vec{r}_N, s_N)$$

(same structure for momentum representation)

Hamiltonian depends on  $\{\vec{r}_i, \vec{p}_i, s_i\}$

$$\hat{H} = \hat{H}(\vec{p}_1, \vec{r}_1, s_1; \dots; \vec{p}_N, \vec{r}_N, s_N)$$

If particles are distinguishable (e.g. different mass, charge, spin, ...) nothing new.

For identical particles new interesting features

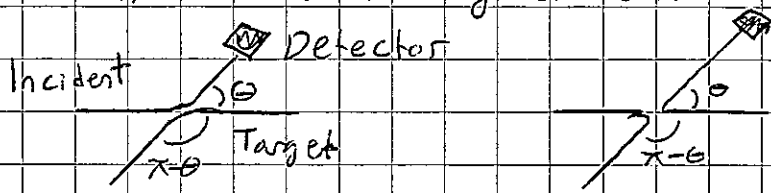
$\hat{H}$  (and other observables  $\rightarrow$  operators) must be symmetric in particle exchange

$$H(\dots; \vec{p}_i, \vec{r}_i, s_i; \dots; \vec{p}_j, \vec{r}_j, s_j; \dots) = H(\dots; \vec{p}_j, \vec{r}_j, s_j; \dots; \vec{p}_i, \vec{r}_i, s_i; \dots)$$

since physical properties of system must not change.

Similar argument holds in classical mechanics, but there particles retain their identity in following well-defined trajectories.

Think of scattering of identical particles:



In classical mechanics, detector rate is the sum of incident and target particles (identical) flying to detector:

$$\frac{d\sigma}{d\Omega} = \frac{d\sigma(\theta)}{d\Omega} + \frac{d\sigma(\pi-\theta)}{d\Omega}$$

incident                      target to det., incident at angle  $\pi - \theta$

In QM, cannot distinguish beam and target particles after collision  $\Rightarrow$  scattering amplitude a superposition

$$\frac{d\sigma}{d\Omega} = |F(\theta) + F(\pi-\theta)|^2 = \frac{d\sigma(\theta)}{d\Omega} + \frac{d\sigma(\pi-\theta)}{d\Omega} + 2\text{Re} F^*(\theta)F(\pi-\theta)$$

$\uparrow$   $\sigma \sim |F|^2$   $\uparrow$  extra term, interference

Define exchange operator  $\hat{P}_{ij}$ , which exchanges two particles  $i, j$

$$\hat{P}_{ij} \psi(\dots; \vec{r}_i, s_i; \dots; \vec{r}_j, s_j; \dots) = \psi(\dots; \vec{r}_j, s_j; \dots; \vec{r}_i, s_i; \dots)$$

Hamiltonian is invariant in this transformation

$$\hat{P}_{ij} \hat{H} \hat{P}_{ij}^{-1} = \hat{H} \quad \Rightarrow \quad [\hat{P}_{ij}, \hat{H}] = 0$$

Clearly  $\hat{P}_{ij}^2 = 1 \quad \Rightarrow \quad \hat{P}_{ij}^{-1} = \hat{P}_{ij}$

Easy to see by integrating

$$\langle \psi | \hat{P}_{ij} \psi \rangle = \langle \psi | \hat{P}_{ij} \psi \rangle^* \Rightarrow \text{Hermitian,}$$

eigenvalues  $\pi_{ij} = \pm 1$

$\pi_{ij}$  constant of motion, eigenstates of  $\hat{H}$  can be chosen as eigenstates of  $\hat{P}_{ij}$ , i.e. symmetric or antisymmetric.

$$\hat{P}_{ij} \psi^S(\dots, i, \dots, j, \dots) = \psi^S(\dots, j, \dots, i, \dots) = +\psi^S(\dots, i, \dots, j, \dots)$$

$$\hat{P}_{ij} \psi^A(\dots, i, \dots, j, \dots) = \psi^A(\dots, j, \dots, i, \dots) = -\psi^A(\dots, i, \dots, j, \dots)$$

Example: two identical particles,  $\psi(1,2)$  any function

$$\psi^S(1,2) = N_S [\psi(1,2) + \psi(2,1)]$$

$$\psi^A(1,2) = N_A [\psi(1,2) - \psi(2,1)]$$

$$N_{S/A} = [2 \pm 2 \operatorname{Re} \langle \psi(1,2) | \psi(2,1) \rangle]^{-1/2}$$

Example: three particles  
completely (anti) symmetric states

$$\psi^S(1,2,3) = N_{3S} [\psi(1,2,3) + \psi(2,3,1) + \psi(3,1,2) + \psi(2,1,3) + \psi(3,2,1) + \psi(1,3,2)]$$

$$\psi^A(1,2,3) = N_{3A} [\psi(1,2,3) + \psi(2,3,1) + \psi(3,1,2) - \psi(2,1,3) - \psi(3,2,1) - \psi(1,3,2)]$$

In addition, there are 4 other linearly independent combinations, with different eigenvalues  $\pi_{ij}$  for different pairs (e.g.  $\pi_{12} = +1$ ,  $\pi_{23} = -1$ ).

These mixed symmetries do not appear in nature!

States of identical particles either fully symmetric or antisymmetric in any pair exchange

symmetric: bosons

$\gamma, \pi, W^\pm, Z, g, \alpha$

integer spin

Bose-Einstein statistics

antisymmetric: Fermions

$q, e^\pm, p, \bar{p}, \nu, {}^3\text{He}$

half-integer spin

Fermi-Dirac statistics

Spin-statistics theorem

VI.2 Fermion system

System of  $N$  noninteracting identical Fermions!

- electrons in atoms
- nucleons in nucleus
- electrons in solids

(interelectron forces approximated by a mean field)

Common potential  $V(\vec{r})$ ,

$$\hat{H} = \sum_{i=1}^N \hat{H}_i(\vec{p}_i, \vec{r}_i, s_i), \quad \hat{H}_i(\vec{p}, \vec{r}, s) = \frac{\vec{p}^2}{2m} + V(\vec{r}, s)$$

one-particle Hamiltonian

Assume eigenstates of  $\hat{H}_i$  known,

$$\hat{H}_i u_k(\vec{r}, s) = E_k u_k(\vec{r}, s)$$

A solution of  $N$ -particle Schrödinger equation would be

$$H\Phi = E\Phi \quad \text{with} \quad \Phi(1, \dots, N) = u_{k_1}(1) u_{k_2}(2) \dots u_{k_N}(N)$$

$$E = E_{k_1} + E_{k_2} + \dots + E_{k_N}$$

but this is not antisymmetric!

Any permutation also solution with same energy

⇒ totally antisymmetric linear combination of permutations

Slater's determinant

$$\Phi(1, \dots, N) = \frac{1}{\sqrt{N!}} \begin{vmatrix} u_{k_1}(1) & u_{k_1}(2) & \dots & u_{k_1}(N) \\ u_{k_2}(1) & u_{k_2}(2) & \dots & u_{k_2}(N) \\ u_{k_3}(1) & & & \vdots \\ \vdots & & & \vdots \\ u_{k_N}(1) & & & u_{k_N}(N) \end{vmatrix}$$

2-particle exchange = exchange two columns, change of sign

Every particle in every state with same probability as others

State known, when we know which one-particle states appear, or are occupied.

If two one-body states are same,  $\det = 0$

$\Rightarrow$  Pauli's exclusion principle: two identical fermions cannot be in same state

Note: determinant normalized correctly, as all 1-body states are orthogonal  $\rightarrow$  permutation terms orthogonal ( $N!$  of them)

Also if two coordinates are same, W.F.  $\rightarrow 0$

$\Rightarrow$  statistical repulsion (even without interaction)

Example: two spin- $\frac{1}{2}$  fermions

same spin: 
$$\frac{1}{\sqrt{2}} \begin{vmatrix} u_1(\vec{r}_1)\uparrow_1 & u_1(\vec{r}_2)\uparrow_2 \\ u_2(\vec{r}_1)\uparrow_1 & u_2(\vec{r}_2)\uparrow_2 \end{vmatrix} = \frac{1}{\sqrt{2}} (u_1(\vec{r}_1)u_2(\vec{r}_2) - u_1(\vec{r}_2)u_2(\vec{r}_1)) \uparrow_1 \uparrow_2$$

|111>

different spin 
$$\frac{1}{\sqrt{2}} \begin{vmatrix} u(\vec{r}_1)\uparrow_1 & u(\vec{r}_2)\uparrow_2 \\ u(\vec{r}_1)\downarrow_1 & u(\vec{r}_2)\downarrow_2 \end{vmatrix} = u(\vec{r}_1)u(\vec{r}_2) \frac{1}{\sqrt{2}} (\uparrow_1 \downarrow_2 - \downarrow_1 \uparrow_2)$$

|100>

symmetric spin  $\Leftrightarrow$  antisymmetric space (triplet)

antisymmetric spin  $\Leftrightarrow$  symmetric space (singlet)

Eg. helium ground state symmetric in space  $\sim$  2 hydrogens  $\Rightarrow$  spins in singlet state  $s=0$

Example 2: expectation value of separation

Distinguishable:  $\psi(1,2) = \psi_a(1)\psi_b(2) \Rightarrow \langle (x_1 - x_2)^2 \rangle_D = \langle x_1^2 \rangle_a + \langle x_2^2 \rangle_b - 2\langle x \rangle_a \langle x \rangle_b$

Identical  $\left. \begin{array}{l} \text{bosons} \\ \text{fermion} \end{array} \right\} \psi = \frac{1}{\sqrt{2}} [\psi_a(1)\psi_b(2) \pm \psi_a(2)\psi_b(1)] \Rightarrow \langle (x_1 - x_2)^2 \rangle_{\pm} = \langle (x_1 - x_2)^2 \rangle_D \mp 2|\langle x \rangle_{ab}|$

If no overlap, no interference

where  $\langle x \rangle_{ab} = \int x \psi_a^*(x) \psi_b(x) dx$

### Inter-particle interactions

$$\hat{H} = \sum_{i=1}^N \hat{H}_1(i) + \sum_{i < j} \hat{V}_2(i,j)$$

~ 2-body operator

In principle  $\Psi$  can be expanded in terms of Slater determinants, but difficult without approximations

Common & useful: mean field approximation

$$\sum_{i < j} \hat{V}_2(i,j) \approx \sum_i \hat{W}(i) ; \hat{H} \rightarrow \frac{\hat{p}^2}{2m} + \hat{V}_1(\vec{r},s) + \hat{W}(\vec{r},s)$$

→ 1-body solutions form Slater determinants, deviations from  $\hat{W}$  residual interactions best treated in perturbation theory

### Shell models:

lowest energy states occupied, each state has one fermion → each spatial state 2, if spin  $-\frac{1}{2}$

atoms:

each shell (nl) can take  $2 \times (2l+1)$  electrons

ns	2	} 8 = period in periodic tables
np	6	

- heavier atoms also d, f states, energies close to other shells: e.g. 3d ~ 4s, 4f ~ 5d ~ 6s
- (- approximate spherically symmetric mean field)

nuclei:

$$|n l s m_l m_s\rangle \rightarrow |n l s j m_j\rangle$$

strong spin-orbit coupling  $\vec{L} \cdot \vec{S} = \frac{1}{2}(J^2 - L^2 - S^2)$

⇒  $m_l, m_s$  not good quantum numbers, need  $\vec{J} = \vec{L} + \vec{S}$  &  $m_j$

energy independent of  $m_j$  ⇒ shells of  $2j+1$  protons +  $2j+1$  neutrons

- classification of shells by  $n_j$ , full shell has  $2(2j+1)$  nucleons

Example: degeneracy pressure

Assume a Fermi gas with all lowest energy states occupied  
remember, that in a box of side length  $L$   $\vec{p} = \frac{2\pi\hbar}{L} (n_x, n_y, n_z)$

$\Rightarrow$  number of states with  $|\vec{p}| \leq p_{max} \equiv p_F$

$$N = \frac{4}{3} \pi p_F^3 \cdot \left(\frac{L}{2\pi\hbar}\right)^3 \cdot 2$$

$$\text{density } g_0 = \frac{N}{V} = \frac{N}{L^3} \stackrel{\text{spin } -1/2}{=} \frac{p_F^3}{3\pi^2 \hbar^3} \equiv \frac{k_F^3}{3\pi^2}$$

$$\rightarrow k_F = (3\pi^2 g_0)^{1/3} \quad \text{Fermi-momentum}$$

$$E_F = \frac{\hbar^2 k_F^2}{2m}$$

Density of states of energy  $E$ :

$$g(E) = \frac{dN}{dE} = \frac{d}{dE} \left( \frac{L^3}{3\pi^2} \left( \frac{2mE}{\hbar^2} \right)^{3/2} \right) = \frac{L^3}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \sqrt{E}$$

$$\begin{aligned} \Rightarrow \text{total energy } E_{tot} &= \int_0^{E_F} dE g(E) E = \frac{L^3}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \int_0^{E_F} dE E^{3/2} \\ &= \frac{L^3}{5\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} E_F^{5/2} = \frac{L^3}{5\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \left[ \frac{\hbar^2}{2m} \left( 3\pi^2 \frac{N}{L^3} \right)^{2/3} \right]^{5/2} \\ &= \frac{\hbar^2}{10m\pi^2} (3\pi^2 N)^{5/3} V^{-2/3} \quad V = L^3 \end{aligned}$$

Pressure:  $dE = -p dV$

$$\Rightarrow p = - \frac{dE}{dV} = \frac{\hbar^2}{15m\pi^2} \left( 3\pi^2 \frac{N}{V} \right)^{5/3} = \frac{\hbar^2}{5m} (3\pi^2)^{2/3} g_0^{5/3}$$

No heat, no interactions, just electrons in ground state

$\Rightarrow$  pressure

This prevents neutron stars and white dwarves from collapsing.